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One- and two-center ETF-integrals of first order in relativistic calculation of NMR parameters

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Abstract

The present work focuses on the analytical and numerical developments of first-order integrals involved in the relativistic calculation of the shielding tensor using exponential-type functions as a basis set of atomic orbitals. For the analytical development, we use the Fourier integral transformation and practical properties of spherical harmonics and the Rayleigh expansion of the plane wavefunctions. The Fourier transforms of the operators were derived in previous work and they are used for analytical development. In both the one- and two-center integrals, Cauchy's residue theorem is used in the final developments of the analytical expressions, which are shown to be accurate to machine precision.

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1. Introduction

Relativistic calculation of nuclear magnetic resonance (NMR) parameters is a subject of significant research [1–14] and still constitutes an important challenge for any of the standard models of quantum chemistry. Calculations involving a magnetic field should preserve gauge independence. This is conveniently accomplished by using gauge including atomic orbitals (GIAO) [15–18].

The main difficulty in the analytical and numerical developments of NMR parameters arises from the operators associated with these parameters, which in the case of the paramagnetic contribution in the relativistic calculations, involve $1/r^5$, where *r* stands for the modulus of the vector \vec{r} that separates two particles. These operators lead to extremely complicated integrals. The finite-perturbation method [19] can be used to compute the NMR parameters [20], but it is well known that numerical differentiation can be very unstable and this is why analytic development has to be used in such calculations.

Magnetic properties of the molecules are particularly sensitive to the quality of the basis sets of atomic orbitals due to many contributing physical phenomena arising from both the vicinity of the nucleus and from the valence region. In ab initio calculations using the linear combination of atomic orbitals-molecular orbitals (LCAO-MO) approximation, molecular orbitals are built from a linear combination of atomic orbitals. Thus, the choice of reliable basis functions is of primary importance. A good atomic orbital basis should satisfy the cusp at the origin [21] and the exponential decay at infinity [22]. The most popular functions used in *ab initio* calculations are Gaussian-type functions (GTFs) [23]. With GTFs, the numerous molecular integrals can be evaluated rather easily. Unfortunately, these GTFs functions fail to satisfy the aforementioned mathematical conditions for atomic-electronic distributions. Consequently, a large number of GTFs have to be used in order to achieve acceptable accuracy, increasing overall computational cost. Exponential-type functions (ETFs) show the same behavior as the exact solutions of atomic or molecular Schrödinger equations satisfying Kato's conditions [24]. Thus, these functions are better suited to represent electron wavefunctions near the nucleus and within a long range. This implies that a smaller number of ETFs is needed for comparable accuracy to the same calculation using GTFs. Unfortunately, molecular multicenter integrals based on ETFs are extremely difficult to evaluate accurately and rapidly. Among ETFs, Slater-type functions (STFs) [25] have a dominating position; this is due to the fact that their analytical expression is very simple. Unfortunately the multi-center integrals over these functions turned out to be extremely difficult to evaluate.

Various studies have focused on the use of *B* functions [26–28], which are analytically more complicated than STFs but have much more appealing properties applicable to multicenter integral problems [27, 29]; in particular, their Fourier transform is exceptionally simple [30]. Note that STFs can be expressed as finite linear combinations of *B* functions [27] and that the basis set of *B* functions is well adapted to the Fourier transform [30–35], which led to analytic expressions for all multi-center molecular integrals over *B* functions.

In the present paper, we derive analytic expressions for one- and two-center integrals of the paramagnetic contribution in the relativistic calculation of the shielding tensor. The analytical development is based on the Fourier transform method. Orthogonality of spherical harmonics and the Rayleigh expansion of the plane wavefunctions are used in the development leading to analytic expressions for the first-order integrals. The obtained analytic expressions involve semi-infinite integrals and their expression is obtained using Cauchy's residue theorem.

2. General definitions and properties

The functions $B_{n,l}^m(\zeta, \vec{r})$ are defined by [27, 28]:

$$B_{n,l}^{m}(\zeta,\vec{r}) = \frac{(\zeta r)^{l}}{2^{n+l}(n+l)!} \hat{k}_{n-\frac{1}{2}}(\zeta r) Y_{l}^{m}(\theta_{\vec{r}},\varphi_{\vec{r}}),$$
(1)

where *n*, *l*, and *m* are the quantum numbers, $\hat{k}_{n-\frac{1}{2}}(z)$ stands for the reduced spherical Bessel function of the second kind defined in [26, 28] and $Y_l^m(\theta, \varphi)$ stands for the surface spherical harmonic [36].

A given function $f(\vec{r})$ and its Fourier transform $\bar{f}(\vec{k})$ are connected by the symmetric relationships:

$$\bar{f}(\vec{k}) = (2\pi)^{-3/2} \int_{\vec{r}} e^{-i\vec{k}\cdot\vec{r}} f(\vec{r}) \,\mathrm{d}\vec{r} \qquad \text{and} \qquad f(\vec{r}) = (2\pi)^{-3/2} \int_{\vec{k}} e^{i\vec{k}\cdot\vec{r}} \,\bar{f}(\vec{k}) \,\mathrm{d}\vec{k}. \tag{2}$$

The Fourier transform $\bar{B}_{n,l}^m(\zeta, \vec{p})$ of $B_{n,l}^m(\zeta, \vec{r})$ is given by [30]

$$\bar{B}_{n,l}^{m}(\zeta,\vec{p}) = \sqrt{\frac{2}{\pi}} \,\zeta^{2n+l-1} \,\frac{(-\mathrm{i}\,|p|)^l}{(\zeta^2+|p|^2)^{n+l+1}} \,Y_l^m(\theta_{\vec{p}},\varphi_{\vec{p}}). \tag{3}$$

The normalized STFs are defined according to the following relationship [25]:

$$\chi_{n,l}^{m}(\zeta,\vec{r}) = \sqrt{\frac{(2\zeta)^{2n+1}}{(2n)!}} r^{n-1} e^{-\zeta r} Y_{l}^{m}(\theta_{\vec{r}},\varphi_{\vec{r}}).$$
(4)

STFs can be expressed as finite linear combinations of *B* functions [27]:

$$\chi_{n,l}^{m}(\zeta,\vec{r}) = \sqrt{\frac{2^{2n+1}\zeta^{3}}{(2n)!}} \sum_{\tilde{n}=\tilde{n}_{\min}}^{\tilde{n}_{\max}} \frac{(-1)^{n-l-\tilde{n}} \ 2^{2\tilde{n}+2l-n} \ (l+\tilde{n})!}{(2\tilde{n}-n+l)! \ (n-l-\tilde{n})!} \ B_{\tilde{n},l}^{m}(\zeta,\vec{r}), \tag{5}$$

where $\tilde{n}_{\text{max}} = n - l$ and $\tilde{n}_{\text{min}} = \frac{n-l}{2}$ if n-l is even, or $\tilde{n}_{\text{min}} = \frac{n-l+1}{2}$ if n-l is odd. Gaunt coefficients are defined by [37, 38]

$$\langle l_1 m_1 | l_2 m_2 | l_3 m_3 \rangle = \int_0^{2\pi} \int_0^{\pi} \left[Y_{l_1}^{m_1}(\theta, \varphi) \right]^* Y_{l_2}^{m_2}(\theta, \varphi) Y_{l_3}^{m_3}(\theta, \varphi) \sin(\theta) \, \mathrm{d}\theta \, \mathrm{d}\varphi.$$
(6)

The Gaunt coefficients linearize the product of two spherical harmonics:

$$\left[Y_{l_1}^{m_1}(\theta,\varphi)\right]^* Y_{l_2}^{m_2}(\theta,\varphi) = \sum_{l=l_{\min},2}^{l_1+l_2} \langle l_2 m_2 | l_1 m_1 | l m_2 - m_1 \rangle Y_l^{m_2-m_1}(\theta,\varphi),$$
(7)

where the subscript $l = l_{\min}$, 2 in the summation symbol implies that the summation index l runs in steps of 2. The constant l_{\min} is given by [38]

$$l_{\min} = \begin{cases} \max(|l_1 - l_2|, |m_2 - m_1|) & \text{if } l_1 + l_2 + \max(|l_1 - l_2|, |m_2 - m_1|) & \text{is even} \\ \max(|l_1 - l_2|, |m_2 - m_1|) + 1 & \text{if } l_1 + l_2 + \max(|l_1 - l_2|, |m_2 - m_1|) & \text{is odd.} \end{cases}$$
(8)

The orthogonality of spherical harmonics is defined by

$$\int_{0}^{2\pi} \int_{0}^{\pi} \left[Y_{l_{1}}^{m_{1}}(\theta,\varphi) \right]^{*} Y_{l_{2}}^{m_{2}}(\theta,\varphi) \sin(\theta) \, \mathrm{d}\theta \, \mathrm{d}\varphi = \delta_{l_{1},l_{2}} \delta_{m_{1},m_{2}}$$
and
$$\int_{0}^{2\pi} \int_{0}^{\pi} Y_{l}^{m}(\theta,\varphi) \sin(\theta) \, \mathrm{d}\theta \, \mathrm{d}\varphi = \delta_{l,0} \delta_{m,0},$$
(9)

where δ_{l_1,l_2} stands for the Kronecker delta function.

The Rayleigh expansion of the plane wavefunctions is given by [39]

$$e^{\pm i\vec{k}\cdot\vec{r}} = 4\pi \sum_{l=0}^{+\infty} \sum_{m=-l}^{l} (\pm i)^{l} \left[Y_{l}^{m}(\theta_{\vec{k}},\varphi_{\vec{k}}) \right]^{*} Y_{l}^{m}(\theta_{\vec{r}},\varphi_{\vec{r}}) j_{l}(k\,r),$$
(10)

where $j_{\lambda}(x)$ stands for the spherical Bessel function of order λ .

The Pochhammer symbol $(\alpha)_n$ is defined by

$$(\alpha)_n = \begin{cases} (\alpha)_n = 1 & \text{if } n = 0, \\ (\alpha)_n = \alpha (\alpha + 1) (\alpha + 2) \cdots (\alpha + n - 1) = \frac{\Gamma(\alpha + n)}{\Gamma(\alpha)} & \text{if } n \leqslant -\alpha, \end{cases}$$
(11)
$$(\alpha)_n = 0 & \text{if } n \geqslant -\alpha + 1, \end{cases}$$

where Γ stands for the Gamma function. For $n \in \mathbb{N}$,

$$\Gamma(n+1) = n!$$
 and $\Gamma\left(n+\frac{1}{2}\right) = \frac{(2n)!}{2^{2n}n!}\sqrt{\pi}.$ (12)

The Cartesian coordinates r_{α} for $\alpha \in \{x, y, z\}$ of a vector \vec{r} can be expressed in terms of spherical polar coordinates and their complex conjugates as

$$r_{\alpha} = r \sum_{m=-1}^{1} c_{\alpha,m} Y_{1}^{m}(\theta_{\vec{r}}, \phi_{\vec{r}}), \qquad (13)$$

where the coefficients $c_{\alpha,m}$ for $\alpha \in \{x, y, z\}$ are given as follows:

$$\begin{cases} c_{x,-1} = \sqrt{\frac{2\pi}{3}}, & c_{y,-1} = i\sqrt{\frac{2\pi}{3}} & \text{and} & c_{z,-1} = 0, \\ c_{x,0} = 0, & c_{y,0} = 0 & \text{and} & c_{z,0} = \sqrt{\frac{4\pi}{3}}, \\ c_{x,1} = -\sqrt{\frac{2\pi}{3}}, & c_{y,1} = i\sqrt{\frac{2\pi}{3}} & \text{and} & c_{z,1} = 0. \end{cases}$$
(14)

3. Analytical development of NMR integrals of the first order

In the presence of an external uniform magnetic field \vec{B}_0 , the electronic non-relativistic Hamiltonian for a system of *n* electrons and *N* nuclei is given by

$$\mathcal{H} = \sum_{i=1}^{n} \left[\frac{1}{2} \vec{p}_i^2 + \sum_{A=1}^{N} \frac{Z_A}{r_{iA}} + \sum_{i< j}^{n} \frac{1}{r_{ij}} \right],\tag{15}$$

where the electron momentum \vec{p}_i is given by

$$\vec{p}_i = [-i\vec{\nabla}_i + e\vec{A}(i)] \qquad \text{where} \quad \vec{A} = \frac{1}{2}(\vec{B}_0 \times \vec{r}_{i0}) + \frac{\mu_0}{4\pi} \sum_N \frac{\vec{\mu}_N \times \vec{r}_{iN}}{r_{iN}^3}, \tag{16}$$

where $\vec{A}(i)$ stands for the vector potential induced by the nuclear moments $\vec{\mu}_N$ and the external uniform magnetic field \vec{B}_0 and μ_0 stands for the dielectric permittivity. Z_A is the atomic number of nucleus A, $\vec{r}_{iA} = \vec{r}_i - \vec{R}_A$, $\vec{r}_{ij} = \vec{r}_i - \vec{r}_j$, \vec{r}_i (resp. \vec{r}_j) represents the vector position of the electron *i* (resp. the electron *j*) and \vec{R}_A is the vector position of the atom *A*.

Heavy atoms require the inclusion of relativistic effects [13] and this gives rise to new parameters, which do not appear in the non-relativistic Hamiltonian. In terms of perturbations with respect to $\mu_{N,\alpha}$ and $B_{0,\beta}$ where α and β stand for Cartesian coordinates ($\alpha, \beta \in (x, y, z)$), the electronic relativistic Hamiltonian is given by

$$\mathcal{H} = \mathcal{H}^{(0)} + \mathcal{H}^{(r)} + \mu_{N,\alpha} \mathcal{H}^{(0,1)}_{\alpha} + B_{0,\beta} \mathcal{H}^{(1,0)}_{\beta} + \mu_{N,\alpha} B_{0,\beta} \mathcal{H}^{(1,1)}_{\alpha\beta} + \cdots,$$
(17)

where $\mathcal{H}^{(0)}$ is the zeroth-order Hamiltonian and $\mathcal{H}^{(r)}$ is the relativistic perturbation term, which is independent of the magnetic perturbations [1, 2].

The first-order perturbation $\mathcal{H}_{\alpha}^{(0,1)}$ of the Hamiltonian with respect to the nuclear moment involves relativistic terms including $\frac{\mu_0}{4\pi} \sum_{j=1}^n 3r_{jN,\beta} \frac{\vec{r}_{jN} \cdot \vec{\sigma}(j)}{r_{jN}^5}$, which correspond to the paramagnetic contribution in the relativistic calculations of the shielding tensor [1, 2]. In this work, we present the analytical and numerical developments of one- and two-center integrals of the first order corresponding to the aforementioned operators.

Let $L_{jN,\alpha}$ be defined by

$$L_{jN,\alpha} = 3r_{jN,\alpha} \left[\frac{\vec{r}_{jN} \cdot \vec{\sigma}_j}{r_{jN}^5} \right],\tag{18}$$

where $\vec{\sigma}_j$ stands for the Pauli spin matrix of the *j*th electron and its Cartesian coordinates are given by

$$\sigma_{j,x} = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}, \qquad \sigma_{j,y} = \begin{pmatrix} 0 & i \\ i & 0 \end{pmatrix} \quad \text{and} \quad \sigma_{j,z} = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}.$$

By developing the scalar product in the expression of $L_{jN,\alpha}$, we obtain

$$L_{jN,\alpha} = 3 r_{jN,\alpha} \frac{\vec{r}_{jN} \cdot \vec{\sigma}_j}{r_{jN}^5} = 3 r_{jN,\alpha} \sum_{\beta} \sigma_{j,\beta} \frac{r_{jN,\beta}}{r_{jN}^5} = -\sum_{\beta} \sigma_{j,\beta} \left[r_{jN,\alpha} \frac{\partial}{\partial r_{jN,\beta}} \left(\frac{1}{r_{jN}^3} \right) \right].$$
(19)

One electron integrals arise from the operator $L_{jN,\alpha}$. These integrals are given by

$$\mathcal{I} = \left\langle B_{n_1,l_1}^{m_1}(\zeta_1, \vec{r}_{jA}) \middle| L_{jN,\alpha} \middle| B_{n_2,l_2}^{m_2}(\zeta_2, \vec{r}_{jB}) \right\rangle_{\vec{r}} \\ = \int_{\vec{r}} \left[B_{n_1,l_1}^{m_1}(\zeta_1, \vec{r}_{jA}) \right]^* 3r_{jN,\alpha} \left[\frac{\vec{r}_{jN} \cdot \vec{\sigma}_j}{r_{jN}^5} \right] B_{n_2,l_2}^{m_2}(\zeta_2, \vec{r}_{jB}) \,\mathrm{d}\vec{r},$$
(20)

where $\vec{r}_{jA} = \vec{r}_j - \overrightarrow{OA} = \vec{r}$, $\vec{r}_{jB} = \vec{r}_j - \overrightarrow{OB} = \vec{r} - \overrightarrow{AB}$ and $\vec{r}_{jN} = \vec{r}_j - \overrightarrow{ON} = \vec{r} - \overrightarrow{AN}$. A, B and N are three arbitrary points of the Euclidian space, while O stands for the origin of the fixed coordinate system.

Using equation (19), the above integrals can be expressed as a finite linear combination of integrals $\mathcal{I}^{(\alpha,\beta)}$ and $\mathcal{I}^{(\alpha)}$. The integrals $\mathcal{I}^{(\alpha,\beta)}$ correspond to the case where α and β represent two different Cartesian coordinates (say $\alpha \neq \beta$) and the integrals $\mathcal{I}^{(\alpha)}$ correspond to the case where α and β represent the same Cartesian coordinate (say $\alpha = \beta$).

The present work concerns the analytical and numerical developments of the integrals $\mathcal{I}^{(\alpha,\beta)}$ ($\alpha \neq \beta$), which are given by

$$\mathcal{I}^{(\alpha,\beta)} = \int_{\vec{r}} \left[B^{m_1}_{n_1,l_1}(\zeta_1,\vec{r}_{jA}) \right]^* \left[r_{jN,\alpha} \frac{\partial}{\partial r_{jN,\beta}} \left(\frac{1}{r_{jN}^3} \right) \right] B^{m_2}_{n_2,l_2}(\zeta_2,\vec{r}_{jB}) \,\mathrm{d}\vec{r} \,. \tag{21}$$

For the analytic development of the above integrals, the Fourier transform method is used. The Fourier transform of the operators was derived in [40] and is given by

$$r_{jN,\alpha} \frac{\partial}{\partial r_{jN,\beta}} \left(\frac{1}{r_{jN}^3} \right) = \sqrt{\frac{2}{\pi}} \frac{k_\alpha k_\beta}{k^2}.$$
(22)

Note that due to the method used to develop the above Fourier transform, it is not valid for the case where α and β represent the same Cartesian coordinate. In this case, the Fourier transform does not exist in the classical sense. The integrals corresponding to $\alpha = \beta$ are a subject of ongoing research.

3.1. One-center integrals

This case occurs when A = B = N. Let us denote by $\vec{r} = \vec{r}_{jN} = \vec{r}_{jA} = \vec{r}_{jB}$. In the following, the one-center integrals will be referred to as $\mathcal{I}_1^{(\alpha,\beta)}$.

Using the Fourier transformation method, we obtain the following expressions for the one-center integrals:

$$\mathcal{I}_{1}^{(\alpha,\beta)} = \frac{1}{2\pi^{2}} \int_{\vec{k}} \frac{k_{\alpha} \, k_{\beta}}{k^{2}} \langle B_{n_{1},l_{1}}^{m_{1}}(\zeta_{1},\vec{r}) \, \big| \, \mathrm{e}^{-\mathrm{i}\vec{k}\cdot\vec{r}} \, \big| B_{n_{2},l_{2}}^{m_{2}}(\zeta_{2},\vec{r}) \big\rangle_{\vec{r}} \, \mathrm{d}\vec{k}. \tag{23}$$

In the term $\mathcal{T}_1 = \langle B_{n_1 l_1}^{m_1}(\zeta_1, \vec{r}) | e^{-i\vec{k}\cdot\vec{r}} | B_{n_2 l_2}^{m_2}(\zeta_2, \vec{r}) \rangle_{\vec{r}}$ involved in the above expression, the two *B* functions are centered at the same point and the radial part of their product is given by

$$\mathcal{T}_{1} = \frac{\sqrt{\pi} \, \zeta_{1}^{l_{1}} \, \zeta_{2}^{l_{2}} \, \zeta_{s}^{l_{1}+l_{2}-1}}{2^{2n_{1}+l_{1}+2n_{2}+l_{2}+1} \, (n_{1}+l_{1})! \, (n_{2}+l_{2})!} \sum_{l=l_{\min},2}^{l_{1}+l_{2}} \frac{(-i)^{l}}{(2\,\zeta_{s})^{l}} \, \langle l_{1}m_{1}|l_{2}m_{2}|lm_{1}-m_{2}\rangle [Y_{l}^{m_{1}-m_{2}}(\theta_{\vec{k}},\varphi_{\vec{k}})]^{*}}{\sum_{\tau=2}^{n_{1}+l_{2}} \sum_{\varsigma=\tau_{1}}^{\tau_{2}} \frac{2^{\tau} \, (2n_{1}-\varsigma-1)! \, (2n_{2}-\tau+\varsigma-1)! \, \zeta_{1}^{\varsigma-1} \zeta_{2}^{\tau-\varsigma-1} \zeta_{s}^{\tau} \, \Gamma(\tau+l_{1}+l_{2}+l+1)}{(\varsigma-1)! \, (n_{1}-\varsigma)! \, (\tau-\varsigma-1)! \, (n_{2}-\tau+\varsigma)! \, \Gamma(l+\frac{3}{2})} \times \sum_{r=0}^{\eta'} \frac{(-1)^{r} \left(\frac{\eta}{2}\right)_{r} \left(\frac{\eta+1}{2}\right)_{r}}{(l+\frac{3}{2})_{r} r! \zeta_{s}^{2r}} \, \frac{k^{l+2r}}{(\zeta_{s}^{2}+k^{2})^{\tau+l_{1}+l_{2}}}, \tag{24}$$

where $\tau_1 = \max(1, \tau - n_2), \tau_2 = \min(n_1, \tau - 1), \zeta_s = \zeta_1 + \zeta_2, \eta = l - \tau - l_1 - l_2 + 1$ and $\eta' = -\frac{\eta}{2}$ if η is even, otherwise $\eta' = -\frac{\eta+1}{2}$. Using equations (13) and (14), we can obtain the following expression for $\frac{k_{\alpha}k_{\beta}}{k^2}$:

$$\frac{k_{\alpha}k_{\beta}}{k^{2}} = \sum_{m_{3}=-1}^{1} \sum_{m_{4}=-1}^{1} (-1)^{m_{3}} c_{\alpha,m_{3}} c_{\beta,m_{4}} \sum_{l'=l'_{\min,2}}^{2} \langle 1m_{4}|1m_{3}|l'm_{4}-m_{3}\rangle Y_{l'}^{m_{4}-m_{3}}(\theta_{\vec{k}},\phi_{\vec{k}}),$$
(25)

where

$$l'_{\min} = \begin{cases} |m_4 - m_3| & \text{if } |m_4 - m_3| & \text{is even,} \\ |m_4 - m_3| + 1 & \text{if } |m_4 - m_3| & \text{is odd.} \end{cases}$$
(26)

The integration of the angular parts of equations (24) and (25) leads to $\int c^{2\pi} \int c^{\pi}$

$$\int_{0}^{\infty} \int_{0}^{\infty} \left[Y_{l}^{m_{1}-m_{2}}(\theta_{\vec{k}},\varphi_{\vec{k}}) \right]^{*} Y_{l'}^{m_{4}-m_{3}}(\theta_{\vec{k}},\phi_{\vec{k}}) \sin(\theta) \, \mathrm{d}\theta \, \mathrm{d}\varphi = \delta_{l,l'} \delta_{m_{1}-m_{2},m_{4}-m_{3}}.$$
(27)

Using the above equation, we obtain the following for $\mathcal{I}_1^{(\alpha,\beta)}$:

$$\begin{aligned} \mathcal{I}_{1}^{(\alpha,\beta)} &= \frac{\zeta_{1}^{l_{1}} \zeta_{2}^{l_{2}} \zeta_{s}^{l_{1}+l_{2}-1}}{\pi^{3/2} 2^{2n_{1}+l_{1}+2n_{2}+l_{2}+2} (n_{1}+l_{1})! (n_{2}+l_{2})!} \sum_{m_{3}=-1}^{1} \sum_{m_{4}=-1}^{1} (-1)^{m_{3}} c_{\alpha,m_{3}} c_{\beta,m_{4}} \\ &\times \sum_{l'=l'\min,2}^{2} \sum_{l=l\min,2}^{l_{1}+l_{2}} \frac{(-i)^{l}}{(2\zeta_{s})^{l}} \langle 1 m_{4} | 1 m_{3} | l' m_{4} - m_{3} \rangle \langle l_{1} m_{1} | l_{2} m_{2} | lm_{1} - m_{2} \rangle \delta_{l,l'} \delta_{m_{1}-m_{2},m_{4}-m_{3}} \\ &\times \sum_{\tau=2}^{n_{1}+n_{2}} \sum_{\varsigma=\tau_{1}}^{\tau_{2}} \frac{2^{\tau} (2n_{1}-\varsigma-1)! (2n_{2}-\tau+\varsigma-1)! \zeta_{1}^{\varsigma-1} \zeta_{2}^{\tau-\varsigma-1} \zeta_{s}^{\tau} \Gamma(\tau+l_{1}+l_{2}+l+1)}{(\varsigma-1)! (n_{1}-\varsigma)! (\tau-\varsigma-1)! (n_{2}-\tau+\varsigma)! \Gamma(l+\frac{3}{2})} \\ &\times \sum_{r=0}^{\eta'} (-1)^{r} \frac{\left(\frac{n}{2}\right)_{r} \left(\frac{n+1}{2}\right)_{r}}{(l+\frac{3}{2})_{r} r! \zeta_{s}^{2r}} \int_{0}^{\infty} \frac{k^{l+2r+2}}{(\zeta_{s}^{2}+k^{2})^{l_{1}+l_{2}+\tau}} dk, \end{aligned}$$

which can be re-written as

$$\mathcal{I}_{1}^{(\alpha,\beta)} = \frac{\zeta_{1}^{l_{1}} \zeta_{2}^{l_{2}} \zeta_{s}^{l_{1}+l_{2}-1}}{\pi^{3/2} 2^{2n_{1}+l_{1}+2n_{2}+l_{2}+2} (n_{1}+l_{1})! (n_{2}+l_{2})!} \sum_{m_{3}=-1}^{1} \sum_{m_{4}=-1}^{1} (-1)^{m_{3}} c_{\alpha,m_{3}} c_{\beta,m_{4}} \\
\times \sum_{l=l_{\min},2}^{2} \frac{(-i)^{l}}{(2\zeta_{s})^{l}} \langle 1 m_{4} | 1 m_{3} | l m_{4} - m_{3} \rangle \langle l_{1} m_{1} | l_{2} m_{2} | l m_{1} - m_{2} \rangle \delta_{m_{1}-m_{2},m_{4}-m_{3}} \\
\times \sum_{\tau=2}^{n_{1}+n_{2}} \sum_{\varsigma=\tau_{1}}^{\tau_{2}} \frac{2^{\tau} (2n_{1}-\varsigma-1)! (2n_{2}-\tau+\varsigma-1)! \zeta_{1}^{\varsigma-1} \zeta_{2}^{\tau-\varsigma-1} \zeta_{s}^{\tau} \Gamma(\tau+l_{1}+l_{2}+l+1)}{(\varsigma-1)! (n_{1}-\varsigma)! (\tau-\varsigma-1)! (n_{2}-\tau+\varsigma)! \Gamma(l+\frac{3}{2})} \\
\times \sum_{r=0}^{\eta'} (-1)^{r} \frac{\left(\frac{\eta}{2}\right)_{r} \left(\frac{\eta+1}{2}\right)_{r}}{\left(l+\frac{3}{2}\right)_{r} r! \zeta_{s}^{2r}} \int_{0}^{\infty} \frac{k^{l+2r+2}}{\left(\zeta_{s}^{2}+k^{2}\right)^{l_{1}+l_{2}+\tau}} dk.$$
(29)

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The last difficulty in the development of $\mathcal{I}_1^{(\alpha,\beta)}$ lies in the semi-infinite integrals involved in its analytical expression (29). We approach this problem by considering the integral

$$\int_0^\infty \frac{k^{l+2r+2}}{\zeta_s^2 + k^2} \,\mathrm{d}k,\tag{30}$$

with the even integrand

$$f(z) = \frac{z^{l+2r+2}}{\zeta_z^2 + z^2}, \qquad z = k + iy.$$
(31)

By considering a positively oriented circular contour above the real axis with a radius $R > \zeta_s$ joined at its two ends by the line along the real axis, and by applying Cauchy's residue theorem and taking the limit as $R \to \infty$, we can write

$$2\int_0^\infty f(k)\,\mathrm{d}k = 2\pi \mathrm{i} \operatorname{Res}_{z=i\zeta_s} f(z). \tag{32}$$

By developing further, we obtain the formula

$$\int_0^\infty \frac{k^{l+2r+2}}{\zeta_s^2 + k^2} \, \mathrm{d}k = \frac{\pi \, i^{l+2r+2}}{2} \zeta_s^{l+2r+1}.$$
(33)

By applying $\frac{1}{(-2)^{\mu-1}\Gamma(\mu)} \left(\frac{\partial}{\zeta_s \partial \zeta_s}\right)^{\mu-1}$, for $\mu = l_1 + l_2 + \tau$ to both sides of the equation and simplifying, we obtain

$$\tilde{\mathcal{I}}_{1}^{(\alpha,\beta)} = \int_{0}^{\infty} \frac{k^{l+2r+2}}{\left(\zeta_{s}^{2}+k^{2}\right)^{l_{1}+l_{2}+\tau}} \, \mathrm{d}k = \frac{\pi \, i^{l+2r+2}}{2} \frac{\left(-r-\frac{l+1}{2}\right)_{l_{1}+l_{2}+\tau-1}}{\Gamma(l_{1}+l_{2}+\tau)} \zeta_{s}^{l+2r+3-2l_{1}-2l_{2}-2\tau}.$$
(34)

Therefore, by simplifying the terms, we finally obtain

$$\mathcal{I}_{1}^{(\alpha,\beta)} = -\frac{\zeta_{1}^{l_{1}}\zeta_{2}^{l_{2}}\zeta_{s}^{2-l_{1}-l_{2}}}{\sqrt{\pi} 2^{2n_{1}+l_{1}+2n_{2}+l_{2}+3} (n_{1}+l_{1})! (n_{2}+l_{2})!} \sum_{m_{3}=-1}^{1} \sum_{m_{4}=-1}^{1} (-1)^{m_{3}} c_{\alpha,m_{3}} c_{\beta,m_{4}} \\
\times \sum_{l=l_{\min},2}^{2} 2^{-l} \langle 1 m_{4} | 1 m_{3} | l m_{4} - m_{3} \rangle \langle l_{1} m_{1} | l_{2} m_{2} | l m_{1} - m_{2} \rangle \delta_{m_{1}-m_{2},m_{4}-m_{3}} \\
\times \sum_{\tau=2}^{n_{1}+n_{2}} \sum_{\varsigma=\tau_{1}}^{\tau_{2}} \frac{2^{\tau} \zeta_{1}^{\varsigma-1} \zeta_{2}^{\tau-\varsigma-1}}{\zeta_{s}^{\tau}} \frac{(2n_{1}-\varsigma-1)! (2n_{2}-\tau+\varsigma-1)! (\tau+l_{1}+l_{2})_{l+1}}{(\varsigma-1)! (n_{1}-\varsigma)! (\tau-\varsigma-1)! (n_{2}-\tau+\varsigma)! \Gamma(l+\frac{3}{2})} \\
\times \sum_{r=0}^{\eta'} \frac{(\frac{n}{2})_{r} (\frac{\eta+1}{2})_{r} (-r-\frac{l+1}{2})_{l_{1}+l_{2}+\tau-1}}{(l+\frac{3}{2})_{r} r!}.$$
(35)

Note that expression (34) obtained using Cauchy's residue theorem implies many simplifications in the overall formula for $\mathcal{I}_1^{(\alpha,\beta)}$ in equation (29) ultimately leading to (35).

3.2. Two-center integrals

This case occurs when $A = B \neq N$. Let us denote by $\vec{r} = \vec{r}_{jA} = \vec{r}_{jB}$ and $\vec{R} = \overrightarrow{AN}$. In the following, the two-center integrals will be referred to as $\mathcal{I}_2^{(\alpha,\beta)}$.

Using the Fourier transformation method, we obtain the following expressions for the two-center integrals:

$$\mathcal{I}_{2}^{(\alpha,\beta)} = \frac{1}{2\pi^{2}} \int_{\vec{k}} \frac{k_{\alpha} \, k_{\beta}}{k^{2}} \, \mathrm{e}^{\mathrm{i}\vec{k}\cdot\vec{R}} \big\langle B_{n_{1},l_{1}}^{m_{1}}(\zeta_{1},\vec{r}) \big| \, \mathrm{e}^{-\mathrm{i}\vec{k}\cdot\vec{r}} \big| B_{n_{2},l_{2}}^{m_{2}}(\zeta_{2},\vec{r}) \big\rangle_{\vec{r}} \, \mathrm{d}\vec{k}.$$
(36)

Following the same development as for the one-center integrals, we obtain

$$\begin{aligned} \mathcal{I}_{2}^{(\alpha,\beta)} &= \frac{1}{2\pi^{2}} \int_{\vec{k}} 4\pi \sum_{\lambda=0}^{+\infty} \sum_{m=-\lambda}^{\lambda} i^{\lambda} \left[Y_{\lambda}^{m}(\theta_{\vec{k}},\varphi_{\vec{k}}) \right]^{*} Y_{\lambda}^{m}(\theta_{\vec{k}},\varphi_{\vec{k}}) j_{\lambda}(R\,k) \\ &\times \sum_{m_{3}=-1}^{1} \sum_{m_{4}=-1}^{1} (-1)^{m_{3}} c_{\alpha,m_{3}} c_{\beta,m_{4}} \sum_{l'=l'_{\min,2}}^{2} \langle 1\,m_{4}|1\,m_{3}|l'\,m_{4}-m_{3} \rangle Y_{l'}^{m_{4}-m_{3}}(\theta_{\vec{k}},\phi_{\vec{k}}) \\ &\times \frac{\sqrt{\pi} \zeta_{1}^{l_{1}} \zeta_{2}^{l_{2}} \zeta_{s}^{l_{1}+l_{2}-1}}{2^{2n_{1}+l_{1}+2n_{2}+l_{2}+1} (n_{1}+l_{1})! (n_{2}+l_{2})!} \sum_{l=l_{\min},2}^{l_{1}+l_{2}} \\ &\times \frac{(-i)^{l}}{(2\,\zeta_{s})^{l}} \langle l_{1}m_{1}|l_{2}m_{2}|lm_{1}-m_{2} \rangle \left[Y_{l}^{m_{1}-m_{2}}(\theta_{\vec{k}},\varphi_{\vec{k}}) \right]^{*} \sum_{\tau=2}^{n_{1}+n_{2}} \sum_{\varsigma=\tau_{1}}^{\tau_{2}} \\ &\times \frac{2^{\tau} (2n_{1}-\varsigma-1)! (2n_{2}-\tau+\varsigma-1)! \zeta_{1}^{\varsigma-1} \zeta_{2}^{\tau-1} \zeta_{s}^{\tau} \Gamma(\tau+l_{1}+l_{2}+l+1)}{(\varsigma-1)! (n_{1}-\varsigma)! (\tau-\varsigma-1)! (n_{2}-\tau+\varsigma)! \Gamma(l+\frac{3}{2})} \\ &\times \sum_{r=0}^{\eta'} \frac{(-1)^{r} \left(\frac{\eta}{2}\right)_{r} \left(\frac{\eta+1}{2}\right)_{r}}{(l+\frac{3}{2})_{r} r! \zeta_{s}^{\varsigma,r}} \frac{k^{l+2r}}{(\zeta_{s}^{2}+k^{2})^{\tau+l_{1}+l_{2}}} \, d\vec{k}. \end{aligned}$$

Using equation (7) to linearize the product of the two spherical harmonics in the above equation, we obtain

$$\begin{aligned} \mathcal{I}_{2}^{(\alpha,\beta)} &= \frac{1}{2\pi^{2}} \int_{\vec{k}} 4\pi \sum_{\lambda=0}^{+\infty} \sum_{m=-\lambda}^{\lambda} i^{\lambda} \left[Y_{\lambda}^{m}(\theta_{\vec{k}},\varphi_{\vec{k}}) \right]^{*} Y_{\lambda}^{m}(\theta_{\vec{k}},\varphi_{\vec{k}}) j_{\lambda}(R\,k) \\ &\times \sum_{m_{3}=-1}^{1} \sum_{m_{4}=-1}^{1} (-1)^{m_{3}} c_{\alpha,m_{3}} c_{\beta,m_{4}} \sum_{l'=l'_{\min,2}}^{2} \langle 1\,m_{4}|1\,m_{3}|l'\,m_{4}-m_{3} \rangle \\ &\times \frac{\sqrt{\pi} \zeta_{1}^{l_{1}} \zeta_{2}^{l_{2}} \zeta_{s}^{l_{1}+l_{2}-1}}{2^{2n_{1}+l_{1}+2n_{2}+l_{2}+1} (n_{1}+l_{1})! (n_{2}+l_{2})!} \sum_{l=l_{\min,2}}^{l_{1}+l_{2}} \frac{(-i)^{l}}{(2\,\zeta_{s})^{l}} \langle l_{1}m_{1}|l_{2}m_{2}|lm_{1}-m_{2} \rangle \\ &\times \sum_{l'=l''_{\min,2}}^{l+l'} \langle l'm_{4}-m_{3}|lm_{1}-m_{2}|l''m_{4}-m_{3}-m_{1}+m_{2} \rangle Y_{l''}^{m_{4}-m_{3}-m_{1}+m_{2}} (\theta_{\vec{k}},\phi_{\vec{k}}) \\ &\times \sum_{\tau=2}^{n_{1}+n_{2}} \sum_{\varsigma=\tau_{1}}^{\tau_{2}} \frac{2^{\tau} (2n_{1}-\varsigma-1)! (2n_{2}-\tau+\varsigma-1)! \zeta_{1}^{\varsigma-1} \zeta_{2}^{\tau-\varsigma-1} \zeta_{s}^{\tau} \Gamma(\tau+l_{1}+l_{2}+l+1)}{(\varsigma-1)! (n_{1}-\varsigma)! (\tau-\varsigma-1)! (n_{2}-\tau+\varsigma)! \Gamma(l+\frac{3}{2})} \\ &\times \sum_{r=0}^{\eta'} \frac{(-1)^{r} \left(\frac{\eta}{2}\right)_{r} \left(\frac{\eta+1}{2}\right)_{r}}{(l+\frac{3}{2})_{r} r! \zeta_{s}^{2r}} \frac{k^{l+2r}}{(\zeta_{s}^{2}+k^{2})^{\tau+l_{1}+l_{2}}} \, d\vec{k}. \end{aligned}$$

The integration of the angular parts $[Y_{\lambda}^{m}(\theta_{\bar{k}},\varphi_{\bar{k}})]^*$ and $Y_{l''}^{m_4-m_3-m_1+m_2}(\theta_{\bar{k}},\phi_{\bar{k}})$ involved in the above equation along with the orthogonality relation of spherical harmonics (9), lead to

$$\mathcal{I}_{2}^{(\alpha,\beta)} = \frac{\zeta_{1}^{l_{1}} \zeta_{2}^{l_{2}} \zeta_{s}^{l_{1}+l_{2}-1}}{\sqrt{\pi} 2^{2n_{1}+l_{1}+2n_{2}+l_{2}} (n_{1}+l_{1})! (n_{2}+l_{2})!} \sum_{l=l_{\min,2}}^{l_{1}+l_{2}} \frac{(-i)^{l}}{(2\zeta_{s})^{l}} \langle l_{1}m_{1}|l_{2}m_{2}|lm_{1}-m_{2}\rangle$$

$$\times \sum_{m_{3}=-1}^{1} \sum_{m_{4}=-1}^{1} (-1)^{m_{3}} c_{\alpha,m_{3}} c_{\beta,m_{4}} \sum_{l'=l'_{\min,2}}^{2} \langle 1 m_{4}|1 m_{3}|l' m_{4}-m_{3}\rangle$$

$$\times \sum_{\lambda=\lambda_{\min,2}}^{l+l} i^{\lambda} \langle l'm_{4} - m_{3}|lm_{1} - m_{2}|\lambda m_{4} - m_{3} - m_{1} + m_{2} \rangle Y_{\lambda}^{m_{4} - m_{3} - m_{1} + m_{2}} \langle \theta_{\vec{R}}, \varphi_{\vec{R}} \rangle$$

$$\times \sum_{\tau=2}^{n_{1}+n_{2}} \sum_{\varsigma=\tau_{1}}^{\tau_{2}} \frac{2^{\tau} (2n_{1} - \varsigma - 1)! (2n_{2} - \tau + \varsigma - 1)! \zeta_{1}^{\varsigma-1} \zeta_{2}^{\tau-\varsigma-1} \zeta_{s}^{\tau} \Gamma(\tau + l_{1} + l_{2} + l + 1)}{(\varsigma - 1)! (n_{1} - \varsigma)! (\tau - \varsigma - 1)! (n_{2} - \tau + \varsigma)! \Gamma(l + \frac{3}{2})}$$

$$\times \sum_{r=0}^{\eta'} \frac{(-1)^{r} (\frac{\eta}{2})_{r} (\frac{\eta+1}{2})_{r}}{(l + \frac{3}{2})_{r} r! \zeta_{s}^{2r}} \int_{0}^{\infty} \frac{k^{l+2r+2}}{(\zeta_{s}^{2} + k^{2})^{l_{1}+l_{2}+\tau}} j_{\lambda}(R k) dk.$$

$$(39)$$

Again, the last difficulty in the development of $\mathcal{I}_2^{(\alpha,\beta)}$ lies in the semi-infinite integrals involved in its analytical expression (39). By proceeding similarly as above, an analytical expression was developed for the integral in [30]. This expression is

$$\int_{0}^{\infty} \frac{k^{\lambda+2n_{k}+2}}{(k^{2}+\alpha^{2})^{\mu+1}} j_{\lambda}(v\,k) \,\mathrm{d}k = \frac{\pi \,(-1)^{n_{k}} \,\alpha^{2n_{k}}}{2^{\mu+1} \,\Gamma(\mu+1) \,v^{\lambda-2\mu+1}} \sum_{s=0}^{n_{k}} \binom{n_{k}}{s} \frac{2^{s} \,(-\mu)_{s}}{(v\,\alpha)^{2s}} \,\hat{k}_{\lambda-\mu+s+\frac{1}{2}}(v\,\alpha).$$
(40)

Re-parameterizing for the integrals in $\mathcal{I}_2^{(\alpha,\beta)}$, we obtain the expression

$$\tilde{\mathcal{I}}_{2}^{(\alpha,\beta)} = \int_{0}^{\infty} \frac{k^{l+2r+2}}{(\zeta_{s}^{2}+k^{2})^{l_{1}+l_{2}+\tau}} j_{\lambda}(R\,k) \, \mathrm{d}k = \frac{\pi \, i^{l+2r-\lambda} \, \zeta_{s}^{l+2r-\lambda}}{2^{l_{1}+l_{2}+\tau} \, \Gamma(l_{1}+l_{2}+\tau) \, R^{\lambda-2l_{1}-2l_{2}-2\tau+3}} \\ \times \sum_{s=0}^{r+\frac{l-\lambda}{2}} \binom{r+\frac{l-\lambda}{2}}{s} \frac{2^{s} \, (1-l_{1}-l_{2}-\tau)_{s}}{(R\,\zeta_{s})^{2s}} \, \hat{k}_{\lambda-l_{1}-l_{2}-\tau+s+\frac{3}{2}}(R\,\zeta_{s}). \tag{41}$$

Inserting this representation into (39), we simplify and develop that

$$\begin{aligned} \mathcal{I}_{2}^{(\alpha,\beta)} &= \frac{\sqrt{\pi} \zeta_{1}^{l_{1}} \zeta_{2}^{l_{2}} \zeta_{s}^{l_{1}+l_{2}-1} R^{2l_{1}+2l_{2}-3}}{4^{n_{1}+l_{1}+n_{2}+l_{2}} (n_{1}+l_{1})! (n_{2}+l_{2})!} \sum_{l=l_{\min,2}}^{l_{1}+l_{2}} 2^{-l} \langle l_{1}m_{1}|l_{2}m_{2}|lm_{1}-m_{2} \rangle \\ &\times \sum_{m_{3}=-1}^{1} \sum_{m_{4}=-1}^{1} (-1)^{m_{3}} c_{\alpha,m_{3}} c_{\beta,m_{4}} \sum_{l'=l'_{\min,2}}^{2} \langle 1m_{4}|1m_{3}|l'm_{4}-m_{3} \rangle \\ &\times \sum_{\lambda=\lambda_{\min,2}}^{l+l'} (R \zeta_{s})^{-\lambda} \langle l'm_{4}-m_{3}|lm_{1}-m_{2}|\lambda m_{4}-m_{3}-m_{1}+m_{2} \rangle Y_{\lambda}^{m_{4}-m_{3}-m_{1}+m_{2}} (\theta_{\vec{R}},\varphi_{\vec{R}}) \\ &\times \sum_{\tau=2}^{n_{1}+n_{2}} \sum_{\zeta=\tau_{1}}^{\tau_{2}} \frac{R^{2\tau} (2n_{1}-\zeta-1)! (2n_{2}-\tau+\zeta-1)! \zeta_{1}^{\zeta-1} \zeta_{2}^{\tau-\zeta-1} \zeta_{s}^{\tau} (\tau+l_{1}+l_{2})_{l+1}}{(\zeta-1)! (n_{1}-\zeta)! (\tau-\zeta-1)! (n_{2}-\tau+\zeta)! \Gamma(l+\frac{3}{2})} \\ &\times \sum_{r=0}^{\eta'} \frac{\left(\frac{\eta}{2}\right)_{r} \left(\frac{\eta+1}{2}\right)_{r}}{(l+\frac{3}{2})_{r} r!} \sum_{s=0}^{r+l-\lambda} \left(\frac{r+\frac{l-\lambda}{2}}{s}\right) \frac{2^{s} (1-l_{1}-l_{2}-\tau)_{s}}{(R \zeta_{s})^{2s}} \hat{k}_{\lambda-l_{1}-l_{2}-\tau+s+\frac{3}{2}} (R \zeta_{s}). \end{aligned}$$
(42)

4. Numerical discussion

Table 1 shows the evaluation of the integral $\tilde{\mathcal{I}}_{1}^{(\alpha,\beta)}$ and table 2 shows the evaluation of the integral $\tilde{\mathcal{I}}_{2}^{(\alpha,\beta)}$ for physically plausible parameters. The values listed as $\tilde{\mathcal{I}}_{1}^{(\alpha,\beta)}$ are obtained using the right-hand side of expression (34) and $\tilde{\mathcal{I}}_{2}^{(\alpha,\beta)}$ are obtained using the right-hand side of expression (41). The values listed as Maple 11 were obtained using the Maple evalf command applied directly to the semi-infinite integral representations of $\tilde{\mathcal{I}}_{1}^{(\alpha,\beta)}$ and $\tilde{\mathcal{I}}_{2}^{(\alpha,\beta)}$. As it can be

	Table 1. Evaluation of $\tilde{\mathcal{I}}_1^{(\alpha,\beta)}$ of equation (34).											
l	r	l_1	l_2	τ	ζ_s	$ ilde{\mathcal{I}}_1^{(lpha,eta)}$	Maple 11					
0	0	1	0	2	0.1	0.196 349 540 849 3621(3)	0.196 349 540 849 3621(3)					
0	1	2	1	3	0.1	0.184 077 694 546 2769(6)	0.184 077 694 546 2769(6)					
0	0	3	0	4	1.0	0.3221359654559846(-1)	0.322 135 965 455 9847(-1)					
0	1	4	1	5	1.0	0.3427488322931979(-2)	0.3427488322931980(-2)					
2	0	0	1	2	1.0	0.589 048 622 548 0862(0)	0.589 048 622 548 0862(0)					
2	1	1	2	3	1.0	0.1840776945462769(-1)	0.184 077 694 546 2769(-1)					
2	2	0	3	2	10.0	0.4295146206079795(-1)	0.429 514 620 607 9795(-1)					

Table 2. Evaluation of $\tilde{\mathcal{I}}_{2}^{(\alpha,\beta)}$ of equation (41).

l	r	l_1	l_2	τ	ζ_s	λ	R	$ ilde{\mathcal{I}}_2^{(lpha,eta)}$	Maple 11
0	0	1	0	2	0.1	0	0.1	0.196 339 788 577 3828(3)	0.196 339 788 577 3828(3)
0	1	2	1	3	0.1	2	1.0	0.122 514 417 148 3751(3)	0.122 514 417 148 3751(3)
0	0	3	0	4	1.0	0	0.1	0.3219570649253009(-1)	0.321 957 064 925 3011(-1)
0	1	4	1	5	1.0	2	1.0	0.8399796133913563(-4)	0.839 979 613 391 3561(-4)
2	0	0	1	2	1.0	0	1.0	0.144 465 918 723 8652(0)	0.144 465 918 723 8652(0)
2	1	1	2	3	1.0	2	10.0	-0.4271414028323236(-4)	-0.4271414028323236(-4)
2	2	0	3	2	10.0	4	1.0	-0.2228563840864297(-5)	-0.2228563840864297(-5)

Table 3. Evaluation of $\mathcal{I}_1^{(x,y)}$ of equation (35).

n_1	l_1	m_1	ζ_1	n_2	l_2	m_2	ζ_2	$\mathcal{I}_2^{(x,y)}$
2	1	-1	2.0	2	1	-1	1.5	0.844 938 515 752 529(-3)
2	1	0	2.0	2	1	0	1.5	0.197895122059999(-3)
3	2	-1	2.0	2	1	-1	1.5	-0.183961338177373(-4)
3	2	1	2.0	3	2	1	1.5	0.314 124 180 426 832(-4)
3	2	2	2.0	3	2	2	1.5	0.132 485 365 811 557(-3)
4	2	1	2.0	2	1	1	1.5	-0.361569492726092(-4)
4	2	1	2.0	3	2	1	1.5	0.764 186 724 377 223(-4)
4	2	2	2.0	3	2	2	1.5	0.331 525 812 542 680(-3)
4	2	0	2.0	4	2	0	1.5	-0.297315128733290(-4)
4	2	1	2.0	4	2	1	1.5	0.215 564 810 358 993(-3)
4	2	2	2.0	4	2	2	1.5	0.951 453 780 055 960(-3)

seen from tables 1 and 2, the expressions obtained from Cauchy's residue theorem are capable of attaining high accuracy.

In tables 3 and 4, we present values for the integrals $\mathcal{I}_1^{(\alpha,\beta)}$ of equation (35). In table 3, we have $\alpha = x$ and $\beta = y$ and in table 4, we have $\alpha = x$ and $\beta = z$. In tables 5–7, we present values for the integrals $\mathcal{I}_2^{(\alpha,\beta)}$ of equation (42). In table 5, we have $\alpha = x$ and $\beta = y$. In table 6, we have $\alpha = x$ and $\beta = z$ and in table 7, we have $\alpha = y$ and $\beta = z$.

Tab	Table 4. Evaluation of $\mathcal{I}_1^{(x,y)}$ of equation (35).												
n_1	l_1	m_1	ζ_1	n_2	l_2	m_2	ζ_2	$\mathcal{I}_2^{(x,y)}$					
3	2	2	2.0	2	1	1	1.5	0.821 954 698 577 322(-3)					
3	2	1	2.0	3	2	0	1.5	0.115 820 024 850 022(-3)					
3	2	-1	2.0	3	2	-2	1.5	-0.511114095358989(-3)					
3	2	2	2.0	3	2	1	1.5	0.738528227838956(-3)					
4	3	2	2.0	3	2	1	1.5	0.684887168407806(-4)					
4	3	3	2.0	3	2	2	1.5	0.465714740331515(-3)					
4	3	3	2.0	4	3	2	1.5	0.487915417374633(-4)					
5	4	2	2.0	4	3	3	1.5	0.172 192 762 141 693(-3)					
5	4	2	2.0	4	3	1	1.5	0.949043394880123(-3)					
5	4	4	2.0	4	3	3	1.5	0.245308514529775(-2)					
5	4	3	2.0	4	3	2	1.5	0.312 594 481 860 128(-2)					
5	4	4	2.0	4	3	3	1.5	0.476597518466772(-2)					
6	5	2	2.0	5	3	3	1.5	0.258993584962462(-2)					
6	5	3	2.0	5	3	2	1.5	0.147851833602632(-1)					
6	5	5	2.0	5	4	4	1.5	0.209404505361548(-2)					

(x =)

Table 5. Evaluation of $\mathcal{I}_2^{(x,y)}$ of equation (42). $\zeta_1 = 1.0, \zeta_2 = 1.5, \vec{R} = (1.5, 75^\circ, 0^\circ)$ in spherical coordinates.

n_1	l_1	m_1	ζ_1	n_2	l_2	m_2	ζ_2	$\mathcal{I}_2^{(x,y)}$
2	1	1	1.0	2	1	-1	1.5	-0.434843064459248(0)
3	2	1	1.0	2	1	-1	1.5	-0.864738363581576(-2)
3	2	1	1.0	2	1	1	1.5	0.480 104 782 258 312(-1)
3	2	1	1.0	3	2	-1	1.5	-0.163214465381496(2)
3	2	2	1.0	3	2	$^{-2}$	1.5	0.410 633 175 483 623(-1)
3	2	2	1.0	3	2	-1	1.5	-0.110224669992678(-1)
3	2	2	1.0	3	2	1	1.5	0.109871879056160(-1)
4	2	1	1.0	2	1	-1	1.5	-0.111977101298055(-1)
4	2	2	1.0	3	2	$^{-2}$	1.5	0.598 123 871 812 077(-1)
4	2	2	1.0	3	2	-1	1.5	-0.160487351821931(-1)
4	2	2	1.0	3	2	1	1.5	0.160078508905662(-1)
4	2	2	1.0	4	2	$^{-2}$	1.5	0.489 662 959 894 970(0)
4	2	2	1.0	4	2	-1	1.5	-0.131339439243965(0)
4	2	2	1.0	4	2	1	1.5	0.131 082 624 218 578(0)
4	3	2	1.0	4	2	1	1.5	0.135 727 821 381 524(2)

For the numerical evaluation of Gaunt coefficients which occur in the complete expressions of the integrals under consideration, we use the subroutine GAUNT.F developed by Weniger *et al* [38]. The spherical harmonics $Y_l^m(\theta, \varphi)$ are computed using the recurrence formulae presented in [38].

In all tables, the numbers in parentheses represent powers of 10.

$\overline{n_1}$	l_1	m_1	ζ_1	n_2	l_2	m_2	ζ_2	$\mathcal{I}_2^{(x,z)}$
2	1	1	1.0	2	1	-1	1.5	-0.993 394 591 541 964(-3)
3	2	1	1.0	2	1	-1	1.5	-0.321257181375785(-1)
3	2	1	1.0	2	1	1	1.5	0.868 646 553 961 828(-2)
3	2	1	1.0	3	2	-1	1.5	-0.220418082414166(-1)
3	2	2	1.0	3	2	$^{-2}$	1.5	0.393 822 816 223 339(-4)
3	2	2	1.0	3	2	-1	1.5	-0.205475160983857(-1)
3	2	2	1.0	3	2	1	1.5	0.163 419 882 900 283(2)
4	2	1	1.0	2	1	-1	1.5	-0.416483489210977(-1)
4	2	2	1.0	3	2	-2	1.5	0.443153951881734(-4)
4	2	2	1.0	3	2	-1	1.5	-0.299240326968214(-1)
4	2	2	1.0	3	2	1	1.5	0.291 681 304 465 606(2)
4	2	2	1.0	4	2	-2	1.5	0.270 308 611 383 939(-3)
4	2	2	1.0	4	2	-1	1.5	-0.244940259990155(0)
4	2	2	1.0	4	2	1	1.5	0.297 907 579 119 817(3)
4	3	2	1.0	4	2	1	1.5	0.729 667 633 180 134(1)

Table 6. Evaluation of $\mathcal{I}_2^{(x,z)}$ of equation (42). $\zeta_1 = 1.0, \zeta_2 = 1.5, \vec{R} = (1.5, 75^\circ, 0^\circ)$ in spherical coordinates.

Table 7. Evaluation of $\mathcal{I}_2^{(y,z)}$ of equation (42). $\zeta_1 = 1.0, \zeta_2 = 1.5, \vec{R} = (1.5, 75^\circ, 0^\circ)$ in spherical coordinates.

n_1	l_1	m_1	ζ_1	n_2	l_2	m_2	ζ_2	$\mathcal{I}_2^{(y,z)}$
2	1	1	1.0	2	1	-1	1.5	-0.980852389239356(-3)
3	2	1	1.0	2	1	-1	1.5	-0.319537014904337(-1)
3	2	1	1.0	2	1	1	1.5	0.858668805557491(-2)
3	2	1	1.0	3	2	-1	1.5	-0.219369423462985(-1)
3	2	2	1.0	3	2	-2	1.5	0.391 684 472 826 858(-4)
3	2	2	1.0	3	2	-1	1.5	-0.204743121872522(-1)
3	2	2	1.0	3	2	1	1.5	0.163 009 587 528 461(2)
4	2	1	1.0	2	1	-1	1.5	-0.414819029773496(-1)
4	2	2	1.0	3	2	-2	1.5	0.441086792217064(-4)
4	2	2	1.0	3	2	-1	1.5	-0.298416140021138(-1)
4	2	2	1.0	3	2	1	1.5	0.291 083 584 089 616(2)
4	2	2	1.0	4	2	-2	1.5	0.269289153349077(-3)
4	2	2	1.0	4	2	-1	1.5	-0.244437213263127(0)
4	2	2	1.0	4	2	1	1.5	0.297 418 176 922 208(3)
4	3	2	1.0	4	2	1	1.5	0.727 410 965 270 183(1)

5. Conclusion

In this paper we show that the Fourier integral transformation can be applied for the analytical development of integrals of the paramagnetic contribution in the relativistic calculation of the shielding tensor using ETFs as a basis set of atomic orbitals. These analytic expressions obtained for the one- and two-center integrals, involve semi-infinite integrals. Cauchy's residue theorem is used in the final developments of the analytical expressions, which are shown to be accurate to machine precision.

The numerical results obtained with the algorithms described in the present work for integrals of the paramagnetic contribution in the relativistic calculation of the shielding tensor over STFs show that it does not seem impossible to envisage that ETFs or related functions may compete with GTFs in accurate and rapid calculations of the NMR properties of molecules in the near future.

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References

- Vaara J 2007 Theory and computation of nuclear magnetic resonance parameters *Phys. Chem. Chem. Phys.* 9 5399–418
- Fukui H and Baba T 1998 Calculation of nuclear magnetic shieldings: XII. Relativistic no-pair equation J. Chem. Phys. 108 3854–62
- [3] Autschbach J and Ziegler T 2002 Relativistic computation of NMR shieldings and spin-spin coupling constants Encycl. Nucl. Magn. Reson. 9 306–23
- Baba T and Fukui H 2002 Calculation of nuclear magnetic shieldings: XIV. Relativistic mass-velocity corrected perturbation Hamiltonians *Mol. Phys.* 100 623–33
- [5] Hess B A 1986 Relativistic electronic-structure calculations employing a two-component no-pair formalism with external-field projection operators *Phys. Rev.* A 33 3742–8
- [6] Wang F and Gou B 2006 Relativistic energy, fine structure, and hyperfine-structure studies of the high-lying core-excited states ${}^{5}P(n)$ (n = 1-3) and ${}^{5}S^{o}(m)$ (m = 1-3) for the Be-like isoelectronic sequence *At. Data Nucl. Data Tables* **92** 176–85
- [7] Visscher L, Enevoldsen T, Saue T, Jensen H J A and Oddershede J 1999 Full four-component relativistic calculations of NMR shielding and indirect spin-spin coupling tensors in hydrogen halides J. Comput. Chem. 20 1262–73
- [8] Schreckenbach G and Ziegler T 1997 Calculation of NMR shielding tensors based on density functional theory and a scalar relativistic Pauli-type hamiltonian. The application to transition metal complexes *Int. J. Quantum Chem.* 61 899–918
- [9] Quiney H M and Belanzoni P 2002 Relativistic calculation of hyperfine and electron spin resonance parameters in diatomic molecules *Chem. Phys. Lett.* 353 253–8
- [10] Moore E A 1980 Relativistic corrections to the hamiltonian for one and two electron atoms Mol. Phys. 39 745–56
- [11] Van Lenthe E, Baerends E J and Snijders J G 1993 Relativistic regular two-component Hamiltonians J. Chem. Phys. 99 4597–610
- [12] Ishida K 2003 Molecular integrals over the Gauge-including atomic orbitals: II. The Breit–Pauli interaction J. Comput. Chem. 24 1874–90
- [13] Pyykkö P 1988 Relativistic effects in structural chemistry Chem. Rev. 88 563-94
- [14] Helgaker T, Jaszuński M and Ruud K 1999 Ab initio methods for the calculation of NMR shielding and indirect spin-spin coupling constants *Chem. Rev.* 99 293–352
- [15] London F 1937 The quantic theory of inter-atomic currents in aromatic combinations J. de Phys. et le Radium 8 397–409
- [16] Ditchfield R 1974 Self-consistent perturbation-theory of diamagnetism: I. Gauge-invariant LCAO method for NMR chemical-shifts *Mol. Phys.* 27 789–807
- [17] Hinton J F, Wolinski K and Pulay P 1990 Efficient implementation of the Gauge-independent atomic orbital method for NMR chemical shift calculations J. Am. Chem. Soc. 112 8251–60
- [18] Gauss J 1994 GIAO-MBPT(3) and GIAO-sDQ-MBPT(4) calculations of nuclear magnetic shielding constants Chem. Phys. Lett. 229 198–203
- [19] Pople J A, McIver J W Jr and Ostlund N S 1968 Self-consistent perturbation theory: II. Nuclear spin coupling constants J. Chem. Phys. 49 2960–5
- [20] Malkina O L, Salahub D R and Malkin V G 1996 Nuclear magnetic resonance spin-spin coupling constants from density functional theory: problems and results J. Chem. Phys. 105 8793–800
- [21] Kato T 1957 On the eigenfunctions of many-particle systems in quantum mechanics Commun. Pure Appl. Math. 10 151–77

- [22] Agmon S 1985 Bounds on Exponential Decay of Eigenfunctions of Schrödinger Operators (Schrödinger operators) ed S Graffi (Berlin: Springer)
- [23] Boys S F 1950 Electronic wave functions: I. A general method of calculation for the stationary states of any molecular system Proc. R. Soc. Lond. A 200 542–54
- [24] Weatherford C A and Jones H W 1982 ETO Multicenter Molecular Integrals (Dordrecht: Reidel)
- [25] Slater J C 1932 Analytic atomic wave functions *Phys. Rev.* **42** 33–43
- [26] Shavitt I 1963 The Gaussian function in calculation of statistical mechanics and quantum mechanics (Methods in Computational Physics: Quantum Mechanics vol 2) ed B Alder, S Fernbach and M Rotenberg (New York: Academic)
- [27] Filter E and Steinborn E O 1978 Extremely compact formulas for molecular one-electron integrals and Coulomb integrals over Slater-type orbitals *Phys. Rev.* A 18 1–11
- [28] Steinborn E O and Filter E 1975 Translations of fields represented by spherical-harmonics expansions for molecular calculations: III. Translations of reduced Bessel functions, Slater-type s-orbitals, and other functions *Theor. Chim. Acta.* 38 273–81
- [29] Weniger E J and Steinborn E O 1989 Addition theorems for *B* functions and other exponentially declining functions *J. Math. Phys.* 30 774–84
- [30] Weniger E J and Steinborn E O 1983 The Fourier transforms of some exponential-type functions and their relevance to multicenter problems J. Chem. Phys. 78 6121–32
- [31] Prosser F P and Blanchard C H 1962 On the evaluation of two-center integrals J. Chem. Phys. 36 1112
- [32] Geller M 1963 Two-electron, one- and two-center integrals J. Chem. Phys. 39 853–4
- [33] Trivedi H P and Steinborn E O 1983 Fourier transform of a two-center product of exponential-type orbitals. Application to one- and two-electron multicenter integrals *Phys. Rev.* A 27 670–9
- [34] Grotendorst J and Steinborn E O 1988 Numerical evaluation of molecular one- and two-electron multicenter integrals with exponential-type orbitals via the Fourier-transform method *Phys. Rev.* A 38 3857–76
- [35] Weniger E J, Grotendorst J and Steinborn E O 1986 Unified analytical treatment of overlap, two-center nuclear attraction and Coulomb integrals of *B* functions via the Fourier-transform method *Phys. Rev.* A 33 3688–705
- [36] Condon E U and Shortley G H 1935 The Theory of Atomic Spectra (Cambridge: Cambridge University Press)
- [37] Gaunt J A 1929 The triplets of helium Phil. Trans. R. Soc. A 228 151–96
- [38] Weniger E J and Steinborn E O 1982 Programs for the coupling of spherical harmonics Comput. Phys. Commun. 25 149–57
- [39] Weissbluth M 1978 Atoms and Molecules (New York: Academic)
- [40] Safouhi H 2009 Integrals of the paramagnetic contribution in the relativistic calculation of the shielding tensor J. Math. Chem. submitted